

Density fluctuations near the QCD critical point in presence of a magnetic field

Mahfuzur Rahaman

Darjeeling Government College

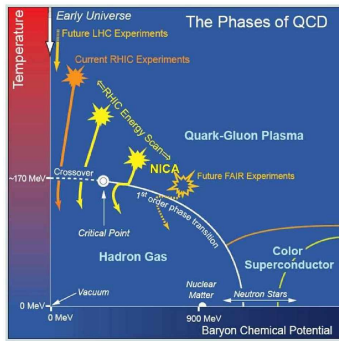
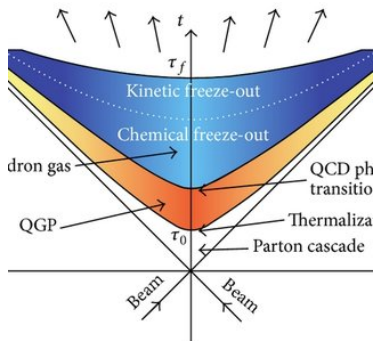
based on arXiv 2306.06905 (MR, Md Hasanujjaman, Golam Sarwar, Abhijit Bhattacharyya, Jan-e Alam)

Meeting on the physics of ALICE, CBM and STAR (MPACS), 29-30 January

- Introduction and motivation**
- Relativistic magneto-hydrodynamics**
- Density fluctuations in QGP & its correlation**
- Results & discussion**
- Summary & outlook**

Introduction

The collisions at LHC, RHIC create a matter, called **quark-gluon plasma**.



- An anticipated critical end point (CEP) at finite μ_B
- We want to investigate if the evolution of the QGP pass through the critical region.

Introduction

- Relativistic Hydrodynamics
- Ideal Hydro: Euler's Equation
- Navier-Stokes **acausal, unstable** Hiscock, Ann.Phys.151,466
- Israel-Stewart Hydrodynamics: **causal, stable** Israel, Stewart Ann.Phys.118,341(1979)

$$T^{\lambda\mu} = \epsilon u^\lambda u^\mu + P \Delta^{\lambda\mu} + q^\lambda u^\mu + q^\mu u^\lambda + \tau^{\lambda\mu}, \quad N^\mu = n u^\mu + n^\mu$$

Conservation equations: $\partial_\mu T^{\mu\lambda} = 0, \quad \partial_\mu N^\mu = 0$

$$\begin{aligned}\Pi &= -\zeta[\partial_\mu u^\mu + \beta_0 D\Pi - \alpha_0 \partial_\mu q^\mu] \\ \pi^{\lambda\mu} &= -2\eta[\Delta^{\lambda\mu\alpha\beta} \partial_\alpha u_\beta + \beta_2 D\pi_{\alpha\beta} - \alpha_1 \partial_\alpha q_\beta] \\ q^\lambda &= \chi T \Delta^{\lambda\mu} \left[-\frac{1}{T} \partial_\mu T - D u_\mu - \beta_1 D q_\mu + \alpha_0 \partial_\mu \Pi + \alpha_1 \partial_\nu \pi_\mu^\nu \right]\end{aligned}$$

Relativistic magneto-hydrodynamics

Non-central collisions \Rightarrow strong magnetic field, $eB \approx 10^{18}$ Gauss! Skokov, et.al. IJMPA 24,5925
For very large σ , the electric field will approach zero.

Constant magnetic field \Rightarrow modify energy-momentum tensor: Gedalin, PRE 51,4901(1995)

$$T_{em}^{\mu\nu} = \frac{B^2}{2} (u^\mu u^\nu - \Delta^{\mu\nu} - 2b^\mu b^\nu), \quad B^\mu B_\mu = -B^2, \quad b^\mu = \frac{B^\mu}{B}$$

For ideal fluid, the total EMT:

$$T_{em}^{\mu\nu}{}_{(0)} = \left(\epsilon + \frac{B^2}{2} \right) u^\mu u^\nu - \left(P + \frac{B^2}{2} \right) \Delta^{\mu\nu} - B^2 b^\mu b^\nu$$

In the presence of polarization (or non-zero magnetization):

$$T_{em}^{\mu\nu}{}_{(0)} = (\epsilon + P - MB) u^\mu u^\nu - (P - MB + \frac{1}{2} B^2) g^{\mu\nu} + (MB - B^2) b^\mu b^\nu$$

Total EMT:

$$T_{tot}^{\mu\nu} = T_{em}^{\mu\nu}{}_{(0)} - \Pi \Delta^{\mu\nu} + q^\mu u^\nu + q^\nu u^\mu + \pi^{\mu\nu}$$

Modified equations:

$$\partial_\mu T_{tot}^{\mu\nu} = 0, \quad \partial_\mu N^\mu = 0$$

Linearised equations

$$0 = \frac{\partial \delta n}{\partial t} + n_0 \vec{\nabla} \cdot \delta \vec{v},$$

$$0 = n_0 \frac{\partial \delta s}{\partial t} + \frac{1}{T_0} \vec{\nabla} \cdot \delta \vec{q} - \frac{1}{n_0 T_0} (B^2 - MB) \frac{\partial \delta n}{\partial t},$$

$$0 = (h_0 + B^2 - MB) \frac{\partial \delta v}{\partial t} + \nabla(\delta P + \delta \Pi) + \frac{\partial \delta q}{\partial t} + \vec{\nabla} \cdot \delta \vec{\pi},$$

$$0 = \delta \Pi + \zeta \left[\vec{\nabla} \cdot \delta \vec{v} + \beta_0 \frac{\partial \delta \Pi}{\partial t} - \alpha_0 \vec{\nabla} \cdot \delta \vec{q} \right],$$

$$0 = \delta \pi^{ij} - \eta \left[\partial^i \delta v^j + \partial^j \delta v^i - \frac{2}{3} g^{ij} \vec{\nabla} \cdot \delta \vec{v} - 2\beta_2 \frac{\partial \delta \pi^{ij}}{\partial t} - \alpha_1 (\partial^i \delta q^j + \partial^j \delta q^i - \frac{2}{3} g^{ij} \vec{\nabla} \cdot \delta \vec{q}) \right],$$

$$0 = \delta q - \kappa T_0 \left[-\frac{\nabla \delta T}{T_0} - \frac{\partial \delta v}{\partial t} - \beta_1 \frac{\partial \delta q}{\partial t} + \alpha_0 \nabla \delta \Pi + \alpha_1 \vec{\nabla} \cdot \delta \vec{\pi} \right].$$

Correlation between Q_i and Q_j is

$$\langle \delta Q_i(\vec{k}, \omega) \delta Q_j(\vec{k}, 0) \rangle$$

The dynamic density factor:

$$S_{nn}(\vec{k}, \omega) = \frac{\langle \delta n(\vec{k}, \omega) \delta n(\vec{k}, 0) \rangle}{\langle \delta n(\vec{k}, 0) \delta n(\vec{k}, 0) \rangle}$$

Dynamic density fluctuation: w/o magnetic field case

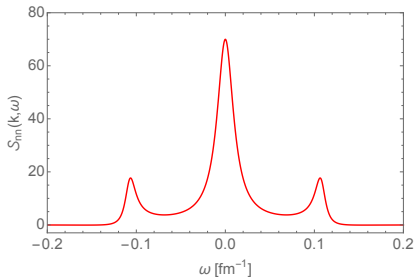


Figure: Away from CEP

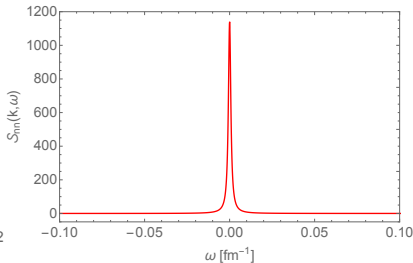


Figure: Near CEP

Dynamic density fluctuation: magnetic field case

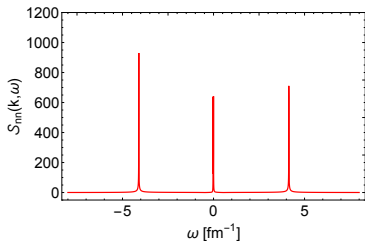


Figure: $eB = m_\pi^2$

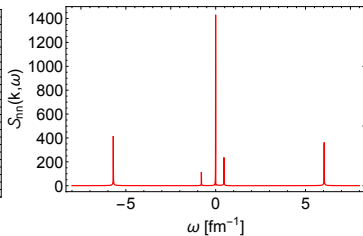
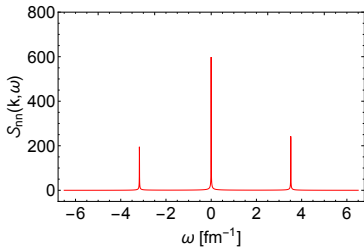


Figure: $eB = 3 m_\pi^2$



Dynamic density fluctuation: CEP and magnetic field

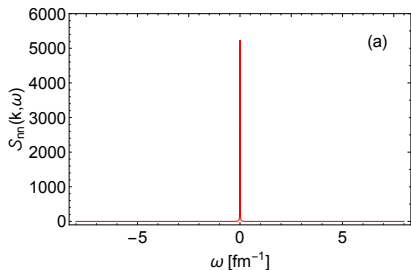


Figure: Near CEP with $B = m_\pi^2 < B_{th}$

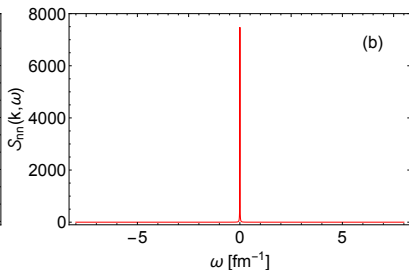


Figure: Near CEP with $B = 3 m_\pi^2 \geq B_{th}$

Summary, Conclusions and Outlook

- Without B , the $S_{nn}(\vec{k}, \omega)$ admits a Rayleigh peak and two Brillouin peaks.
- In B , two extra peaks appearing closer to the R-peak, identified as the B-peaks due to the adiabatic transverse pressure fluctuation.
- The extra B-peaks (near side) exclusively appear beyond a threshold value of the magnetic field.
- The away side and the near side B-peaks are caused by the pressure fluctuation in longitudinal and transverse direction respectively due to pressure difference $P_{\perp} = P - MB$ and longitudinal, $P_L = P$ components.
- This inherent pressure anisotropy will subsequently influence the flow harmonics.
- The $S_{nn}(\vec{k}, \omega)$ for Navier-Stokes reveals near side B-peaks are absent but away side B peaks are present although their position and magnitude differs. It suggests that near side B peaks appear due to coupling between B and the coupling and relaxation coefficients of 2nd order theory.

Thank You